

# Left-Right Symmetry: At the Edges of Phase Space and Beyond <sup>1</sup>

UMass - Amherst

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**elusives**  
neutrinos, dark matter & dark energy physics



**inVisiblesPlus**

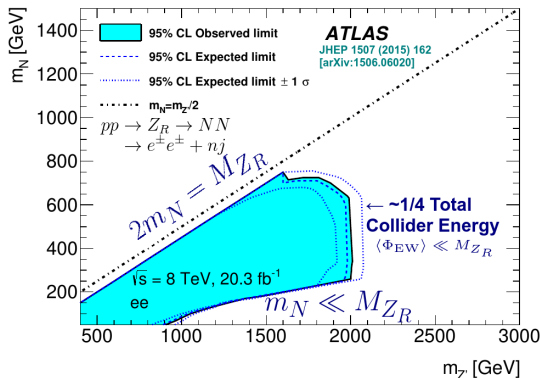
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<sup>1</sup>Based on several works: See slides for Refs. (\*) = IPPP student.

# An Emerging Picture of New Physics Physics

The LHC is operating amazingly!  $\mathcal{L} \sim 40 \text{ fb}^{-1}$  at 13 TeV ( $3\text{-}4 \times \mathcal{L}_{\text{Tevatron}}$ )

**Plotted:** Excluded  $(m_N, M_{Z_R})$  from  $pp \rightarrow Z_R \rightarrow NN$  searches



While no confirmed **BSM** discoveries at colliders, it certainly still possible

- Remaining model space is hierarchical  $\Rightarrow$  extrema of phase space

## Left-Right Symmetry...

When hierarchies are present, often a qualitatively different picture emerges.  $\Rightarrow$  **Quantitatively**, difficult problems become simpler to solve.

- E.g., Effective Field Theory, Hadronization, Classical Mechanics

**Question:** Does collider pheno for neutrino mass models (Seesaws) qualitatively change for hierarchical regions of model parameter space?

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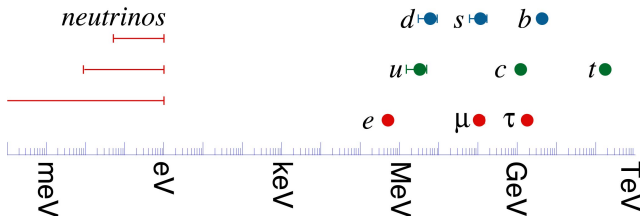
... At the Edges of Phase Space and Beyond:

- 1 Left-Right Symmetric Model Primer
- 2 LRSM at the Edges of Phase Space
- 3 LRSM beyond the Edges of Phase Space
- 4 Redux I: Edges
- 5 Redux II: Beyond

## Motivation for new physics from $\nu$ physics

# Our Motivation

The SM, via the Higgs Mechanism, explains *how* elementary fermions obtain mass, i.e., the  $m_f = y_f \langle \phi \rangle$ , **not** the values of  $m_f$ .



Spanning many orders of magnitudes, the relationship of fermion masses is still a mystery. Two observations:

- 1 Neutrinos have mass (BSM physics! 🏵️!)
- 2 Neutrinos have unusually small mass (new physics? 🏵️?)

# Collider Connection to Neutrino Mass Models (1/1)

Seesaw models predict new Seesaw particles of all shapes, spins, and color:

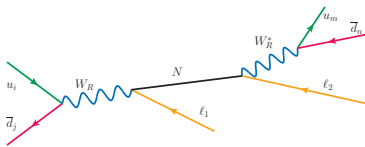
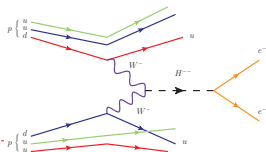
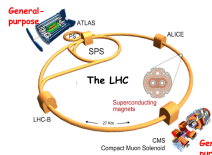
$N$  (Type I),  $T^{0,\pm}$  (Type III),  $Z_{B-L}$ ,  $H_R^{\pm,\pm\pm}$  (Type I+II), ...

Through gauge couplings and mixing, production in  $ee/ep/pp$  collisions

$$\text{DY} : q\bar{q} \rightarrow \gamma^*/Z^* \rightarrow T^+T^- \quad \text{and} \quad q\bar{q}' \rightarrow W_R^\pm \rightarrow N\ell^\pm$$

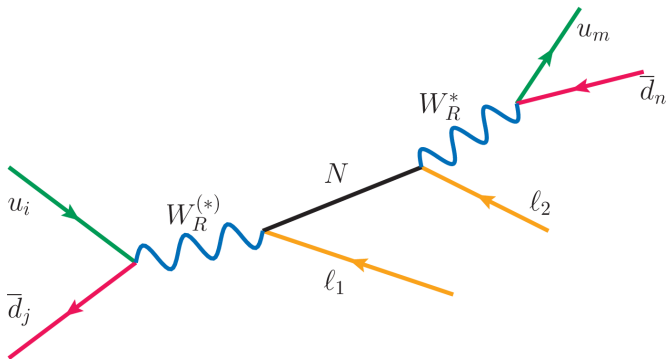
$$\text{VBF} : W^\pm W^\pm \rightarrow H^{\pm\pm}$$

$$\text{GF} : gg \rightarrow h^*/Z^* \rightarrow N\nu_\ell$$



**Identification** of Seesaw partners is then inferred by their decays to SM particles and the associated final-state kinematics

## Left-Right Symmetry at Hadron Colliders





Left-Right Symmetric Models (**LRS**M) postulate that the SM's  $V - A$  structure originates from the spontaneous breakdown of parity symmetry:

$$\mathrm{SU}(3)_c \otimes \mathrm{SU}(2)_L \otimes \underbrace{\mathrm{SU}(2)_R \otimes \mathrm{U}(1)_{B-L}}$$

After scalar  $\Delta_R$  acquires a vev  $v_R \gg v_{SM}$ :  $\hookrightarrow \mathrm{U}(1)_Y$

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With  $N_R$ , all SM fermions can be grouped in  $SU(2)_L$  and  $SU(2)_R$  doublets. Dirac masses generated in (mostly) usual way with  $\Phi$ , i.e.,  $\Delta\mathcal{L} \ni \bar{Q}_L \Phi Q_R$

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Neutrinos obtain LH (RH) Majorana masses from triplet scalar  $\Delta_L$  ( $\Delta_R$ ):

$$m_{\text{light}}^\nu = \underbrace{y_L \langle \Delta_L \rangle}_{\text{Type II}} - \underbrace{\left( y_D y_R^{-1} y_D^T \right) \langle \Phi \rangle^2 \langle \Delta_R \rangle^{-1}}_{\text{Type I a la Type II}} \sim \mathcal{O}(0) + \text{symm.-breaking}$$

**Major pheno:** heavy  $N$ ,  $W'/Z'$  ( $\approx W_R/Z_R$ ), and  $H_i^{\pm\pm}$ ,  $H_j^\pm$ ,  $H_k^0$

$W_R$  coupling to quarks is analogous to SM  $W_{SM}$  couplings:

$$\mathcal{L} = -\frac{g}{\sqrt{2}} W_{R\mu}^- \sum_{q=u,d,\dots} [\bar{d}_j V_{ij}^{CKM'} \gamma^\mu P_R u_i] + \text{H.c.}$$

In **chiral/gauge** basis, couplings to leptons is given by:

$$\mathcal{L} = -\frac{g}{\sqrt{2}} W_{R\mu}^- \sum_{a=1}^3 [\bar{l}^a \gamma^\mu P_R \underbrace{N_R^a}_{\text{Note: } |N_R\rangle = X|v_m\rangle + Y|N_{m'}\rangle}] + \text{H.c.}$$

This is not a practical basis to use.

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In **mass** basis, coupling to leptons can be generically parametrized as<sup>2</sup>:

$$\mathcal{L} = -\frac{g}{\sqrt{2}} W_{R\mu}^- \sum_{\ell=e}^{\tau} \bar{\ell} \gamma^\mu P_R \left[ \underbrace{\sum_{m=1}^3 \underbrace{X_{\ell m}}_{\mathcal{O}(m_\nu/m_N)} \nu_m + \sum_{m'=4}^6 \underbrace{Y_{\ell m'}}_{\mathcal{O}(1)} N_{m'}}_{\text{Note: } |N_R\rangle = X|\nu_m\rangle + Y|N_{m'}\rangle} \right] + \text{H.c.}$$

<sup>2</sup>Atre, Han, Pascoli, Zhang [0901.3589]; Han, Lewis, RR, Si [1211.6447]

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**Benchmark:** Simply consider only the lightest  $N \equiv N_{m'=4}$  and that the  $N$  mass state is aligned with  $\ell = e$  flavor state, i.e.,  $|Y_{eN}| = 1$ .

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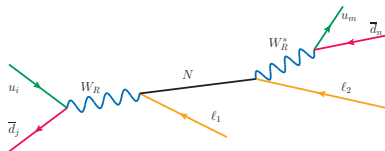
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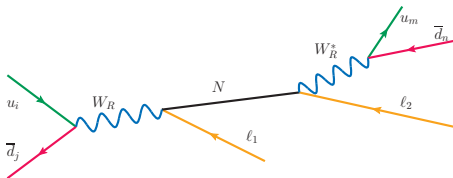
- $W_R \rightarrow Ne$  decay rate is  $\text{BR} \approx 10\%$  for  $M_{W_R} \gg m_N$  (90% to quarks)
- $N \rightarrow W_R^{*\pm} \ell^\mp \rightarrow \ell^\mp qq'/tb$  are dominant decay channels with  $\text{BR} \approx 100\%$



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Hallmark LRSM collider signature is the spectacular same-sign lepton pairs:

$$q\bar{q}' \rightarrow W_R^\pm \rightarrow N \ell_1^\pm \rightarrow \ell_1^\pm \ell_2^\pm q'\bar{q}$$



Proposed by Keung & Senjanovic ('83) and basis for most Seesaw searches

- $W_R^\pm$  is heavy<sup>4</sup>. If kinematically accessible, s-channel  $q\bar{q}' \rightarrow W_R^\pm$  production rate is largest at LHC
- **$L$ -violating process!**  $\Rightarrow$  Majorana nature of  $\nu$  [Black Box Theorem]
- $W_R^* \rightarrow q'\bar{q}$  allows for full reconstruction of kinematics/properties
- High- $p_T$   $\ell^\pm$  without light  $\nu \Rightarrow$  no transverse mom. imbalance (**MET**)

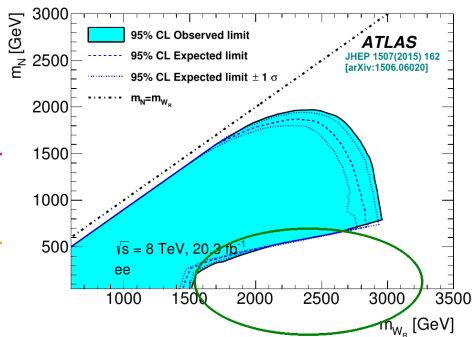
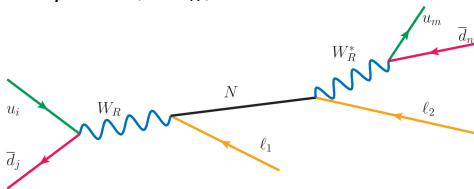
<sup>4</sup>ATLAS [1506.06020; 1512.01530] and CMS [1407.06020; 1512.01224]

# 8 TeV LHC Exclusion with $\mathcal{L} \approx 20 \text{ fb}^{-1}$

LHC expts have performed remarkably!

**Plotted:** excluded  $(m_{N_R}, M_{W_R})$  from searches for resonant  $W_R, N$

**Signature:**  $pp \rightarrow e^\pm e^\pm + nj + X$   
 $+ p_T^\ell \gtrsim \mathcal{O}(M_{W_R}) + \text{no MET}$

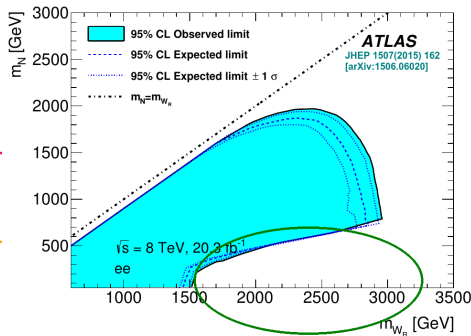
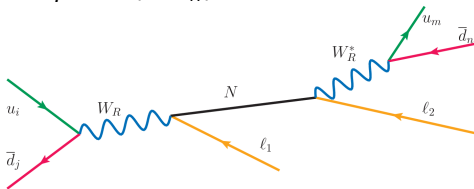


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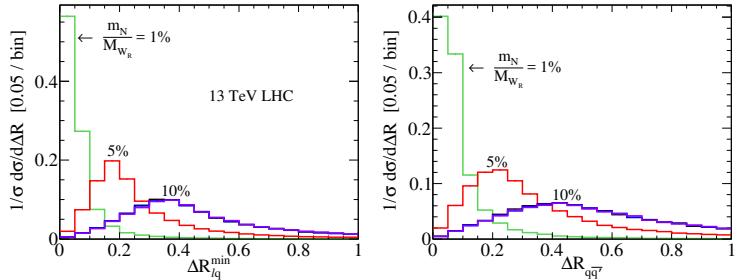


Similar sensitivity to searches for  $pp \rightarrow Z_R \rightarrow NN \rightarrow e^\pm e^\pm + nj + X$

$\Rightarrow$  For both  $W_R$  and  $Z_R$ , loss of sensitivity when  $m_N \ll M_V$

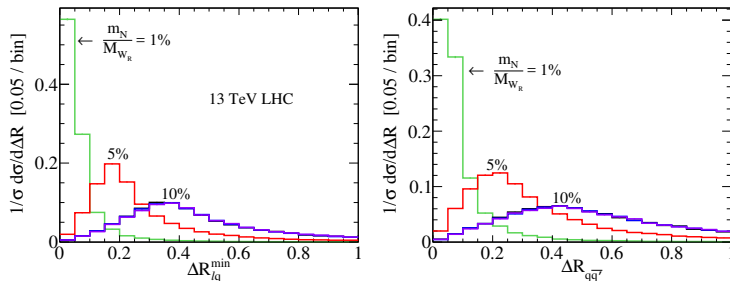
(Lets see what is going on.)

# Failure of Electron ID in $pp \rightarrow W_R \rightarrow \ell^\pm N(\rightarrow \ell^\pm q\bar{q}')$



For a  $1 \rightarrow 2$  process,  $m_{ij}^2 = (p_i + p_j)^2 \approx 2E_i E_j (1 - \cos \theta_{ij}) \approx E_i E_j \theta_{ij}^2$

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$$\Rightarrow \Delta R_{ij} \sim \frac{m_N}{\sqrt{E_i E_j}} \sim \frac{4m_N}{M_{W_R}} \Rightarrow \text{For } \left( \frac{m_N}{M_{W_R}} \right) < 0.1, \Delta R_{\ell X}^{\min} = 0.4 \text{ iso. req. fails}$$

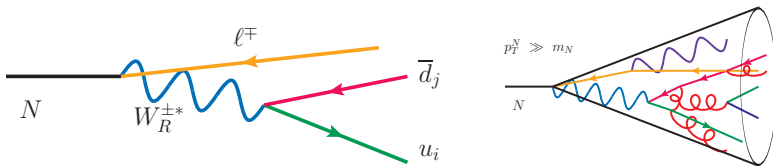
K&S process  $pp \rightarrow \ell^\pm \ell^\pm jj + X$  contains two same-sign charged leptons  
 - S/B power comes from high- $p_T$  leptons without accompanying MET

**Question:** Is it necessary to identify the second lepton or jet multiplicity?

## Neutrino Jets<sup>5</sup> (n):

(i) hadronically decaying, high- $p_T$  heavy neutrinos;

(ii) a fat jet originating from a heavy neutrino



<sup>5</sup>A. Ferrari, et, al, PRD ('00); Mitra, **RR**, Scott\*, Spannowsky, PRD ('16) [1607.03504]; Mattelaer, Mitra, **RR** [1610.08985]

# Jet Structure and Substructure

Consider the Higgs decay  $h \rightarrow b\bar{b}$  with  $p_h = p_j + p_k$  and  $z = E_j/E_h$ .

$$m_h^2 = (p_j + p_k)^2 \approx z(1-z)E_h^2\theta_{jk}^2$$
$$\rightarrow \Delta R_{jk} \sim \theta_{jk} = \frac{m_h}{\sqrt{z(1-z)}E_h} = \frac{2m_h}{E_h}$$

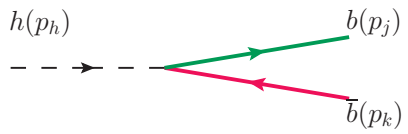
In  $1 \rightarrow 2$  decays we have  $z = (1-z) = 0.5 \implies \frac{1}{\sqrt{z(1-z)}} = 2$ .

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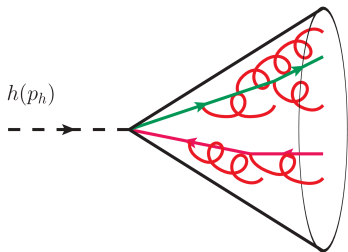


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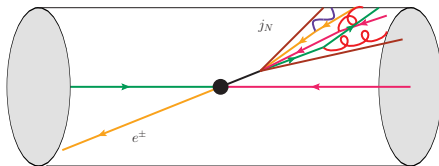


Collimated objects are difficult to resolve.

**Solution:** Instead of treating decay products as individual objects, consider them as a single object, a *Higgs jet*.

# Neutrino Jets in LRSM

Change the scale of our problem: treat  $\ell_2^\pm$  like any other poorly isolated parton bathed in QCD radiation and cluster via a sequential jet algorithm<sup>6</sup>



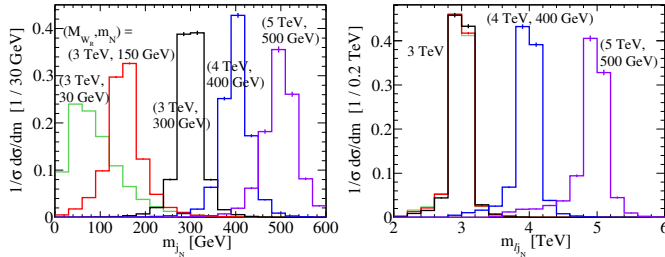
Changing scales *simplifies* the problem, a lot:

For  $m_N \ll M_{W_R}$ , one has a different collider topology:

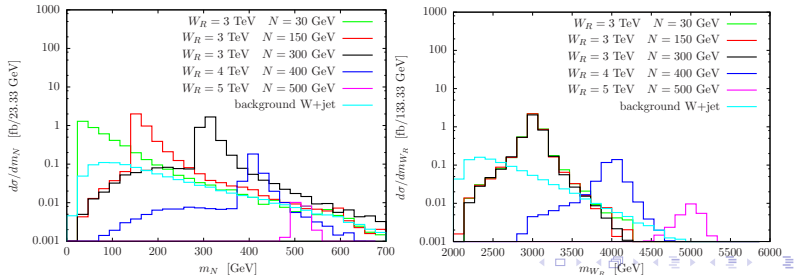
$$pp \rightarrow W_R \rightarrow e^\pm N \rightarrow e^\pm j_{\text{Fat}} \text{ (+ no MET!)}$$

<sup>6</sup>Sequential jet algorithm  $\approx$  definition of collimated, clusters of partons that is meaningful at all orders of perturbation theory, i.e., Infrared-Collinear (IRC)-safe

At parton-level + smearing, expected invariant mass peaks are visible:

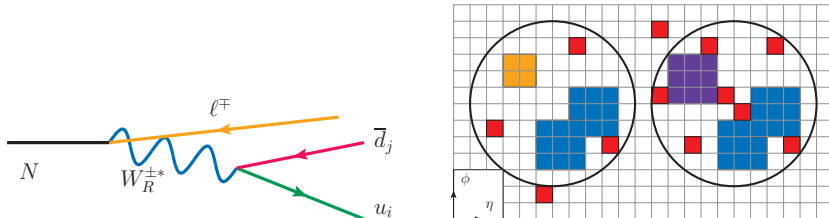


QCD corrections do not change this; oddly, exhibits “ideal” jet behavior  
 With parton shower + P.U. + detector simulation, structures are retained:

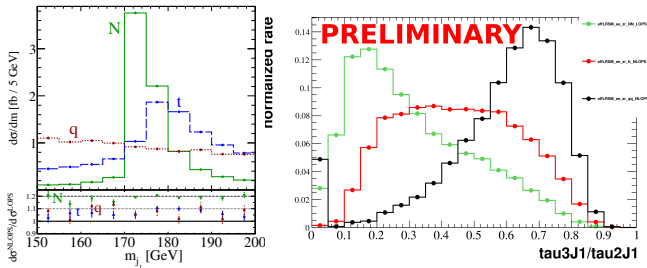


Neutrino jets inherently contain less QCD radiation than top jets

- Prongs within a jet are more likely to be resolved

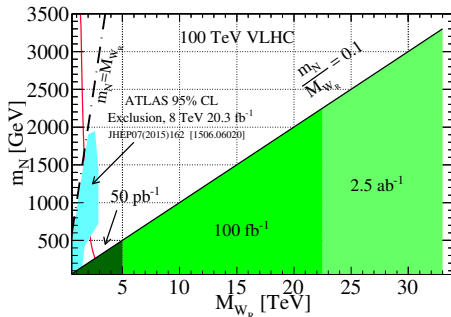
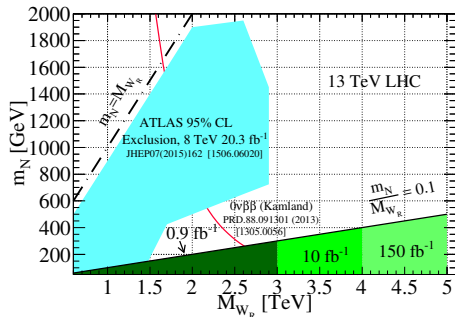


Retain resonant structure and substructure more ideally than top jets:



# Discovery Potential at the Edge of Phase Space

For  $m_N/M_{W_R} \leq 0.1$ , the region where ATLAS/CMS searches breakdown, neutrino jet searches recovers lost sensitivity



**Signature:**  $pp \rightarrow \ell^\pm + j_{\text{Fat}} + X$  [no MET,  $p_T^{\ell,j} \gtrsim 1$  TeV,  $M_{\ell j}$  Cut]

- 13 TeV:  $M_{W_R} \approx 3$  (4) [5] TeV discovery after 10 (100) [2000]  $\text{fb}^{-1}$
- 100 TeV:  $M_{W_R} \approx 15$  (30) TeV discovery after 100  $\text{fb}^{-1}$  (10  $\text{ab}^{-1}$ )

# Left-Right Symmetry Beyond the Edge of Phase Space:

# **Left-Right Symmetry Beyond the Edge of Phase Space:**

## **A pathological but plausible scenario.**

Limits on neutral flavor changing transitions require<sup>7</sup>  $\Delta_R$  sector to be  $\langle \Delta_R \rangle \gtrsim \mathcal{O}(10)$  TeV

What if LR gauge and Yukawa couplings have similar values as in the SM?

- What if  $M_{W_R} \sim g_L \langle \Delta_R \rangle \sim 6.5$  TeV and  $m_N \sim y_{\text{SM}}^\tau \langle \Delta_R \rangle \sim 100$  GeV?

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<sup>7</sup>Bertonlini, et al [1403.7112, + others]; Zhang, et al. [0704.1662; + others]

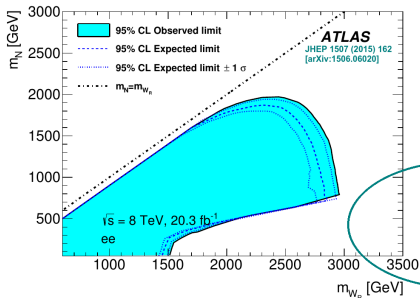
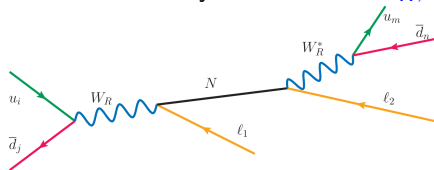


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**Data** may be suggesting EW-scale  $N$   
but kinematically inaccessible  $W_R, Z_R$



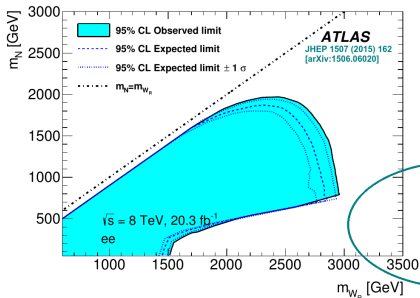
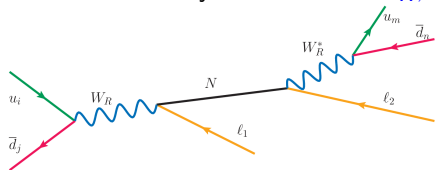
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Data may be suggesting EW-scale  $N$  but kinematically inaccessible  $W_R, Z_R$



Searches follow Keung & Senjanovic ('83), and assume resonant  $W_R, N$

- Zero sensitivity to  $M_{W_R} > 6 - 7$  TeV due to finite data set
- **Naive Question:** is an on-shell  $W_R$  necessary for discovery of  $N$ ?

<sup>7</sup>Bertonlini, et al [1403.7112, + others]; Zhang, et al. [0704.1662; + others]

**Of course**  $pp \rightarrow W_R^* \rightarrow N\ell + X$  can occur via an off-shell mediator.

- Simply LR analog of Fermi contact interaction  $\mathcal{L} = G_F [\bar{N} \gamma^\mu \mathcal{P}] [\bar{\nu} \gamma_\mu \ell]$

Interestingly, in the limit that  $M_{W_R} \gg \sqrt{\hat{s}}$  but  $m_N \lesssim \mathcal{O}(1)$  TeV,  
 $pp \rightarrow N\ell + X$  production in the LRSM and “Type I” are indistinguishable<sup>8</sup>

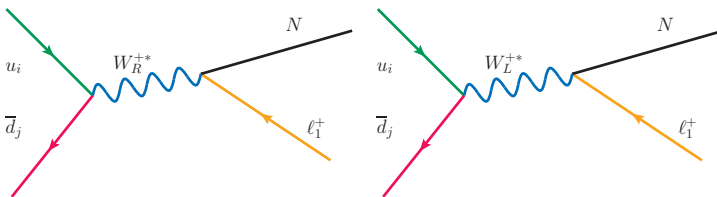
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- Occurs near threshold  $\sqrt{\hat{s}} \sim m_N$  and same  $\ell_1^\pm$  polarization
- Differentiation requires polar and azimuthal polarization measurements of the full  $pp \rightarrow \ell^\pm \ell^\pm + nj + X$  final state

<sup>8</sup>Han, Lewis, **RR**, Si, PRD ('12) [1211.6447]; **RR**, EPJC ('17) [1703.04669]

# $L$ Violation from Beyond the Edges of Phase Space<sup>9</sup>

“**Type I**” searches and projected sensitivities for can be reinterpreted in the context of LRSM in the limit that  $M_{W_R} \sim \sqrt{s} \gg \sqrt{\hat{s}}$

- **Signature:**  $pp \rightarrow \ell^\pm \ell^\pm + nj + X + p_T^\ell \gtrsim \mathcal{O}(m_N) + \text{no MET}$

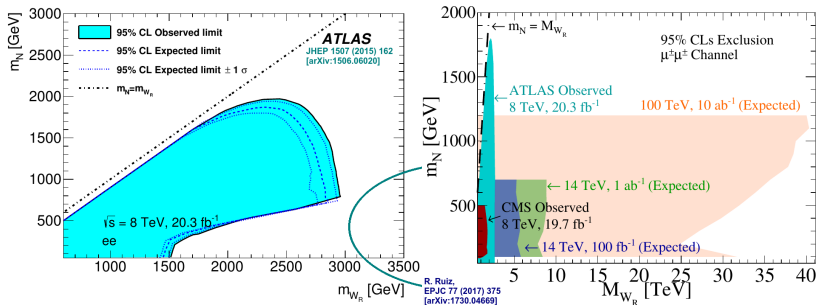
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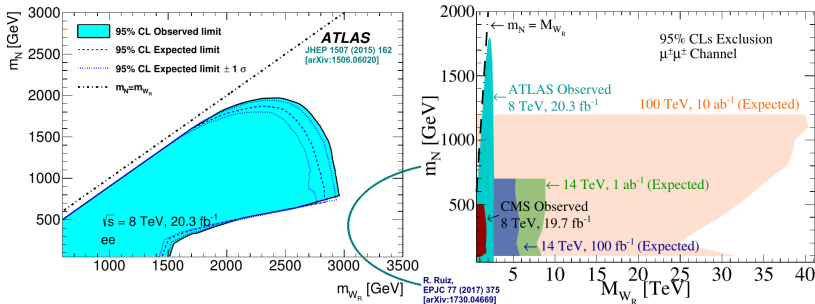


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At 14 (100) TeV with  $\mathcal{L} = 1$  (10) ab<sup>-1</sup>,  $M_{W_R} \lesssim 9$  (40) TeV can be probed

- Caveat: Numbers can be improved with (a) dedicated analysis (not reinterpretation) and (b) knowledge of 100 TeV detector definition

<sup>9</sup>RR, EPJC ('17) [1703.04669]

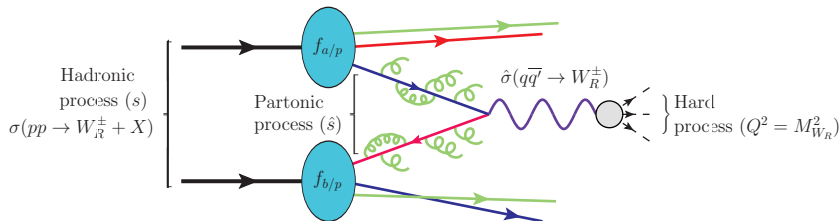
## Redux I: Back to Edges of the LHC Phase Space



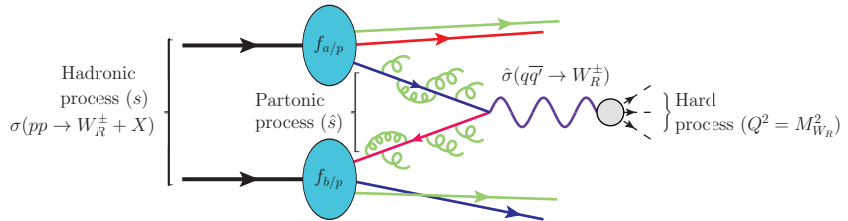
## Redux I: Back to Edges of the LHC Phase Space

Can you see  $M_{W_R} \gtrsim 5 \text{ TeV}$ ?

**Recall:**  $W_R$  production is analogous to  $W_{\text{SM}}$ , except  $M_{W_R} \gtrsim 3 - 5 \text{ TeV}$

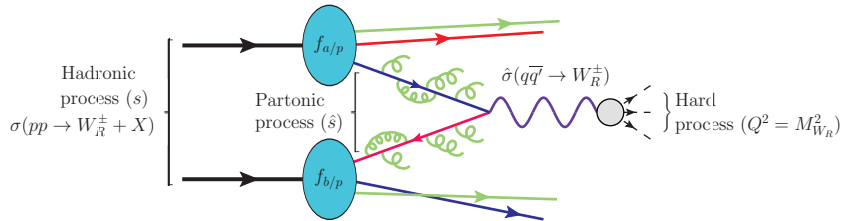


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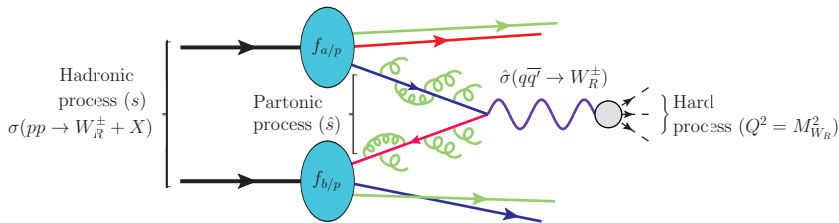
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**Away** from phase space boundaries, QCD corrections are 20-30%.

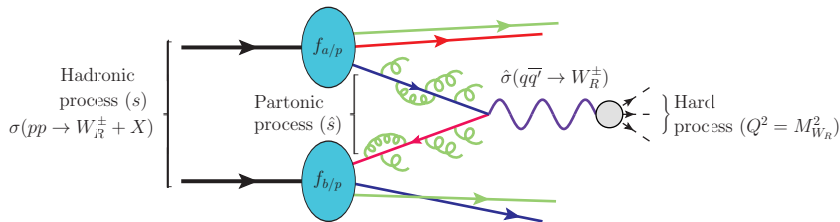
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$$\begin{aligned} \sigma(pp \rightarrow W_R + g) &\sim \int d^{4-2\varepsilon} PS_2 \sim \lambda^{\frac{1-2\varepsilon}{2}} \left( 1, \frac{Q^2=M_{W_R}^2}{\hat{s}}, \frac{k_g^2=0}{\hat{s}} \right) \\ &= \left( 1 - \frac{M_{W_R}^2}{\hat{s}} \right)^{1-2\varepsilon} \sim 2\varepsilon \log \left( 1 - \frac{M_{W_R}^2}{\hat{s}} \right) \end{aligned}$$

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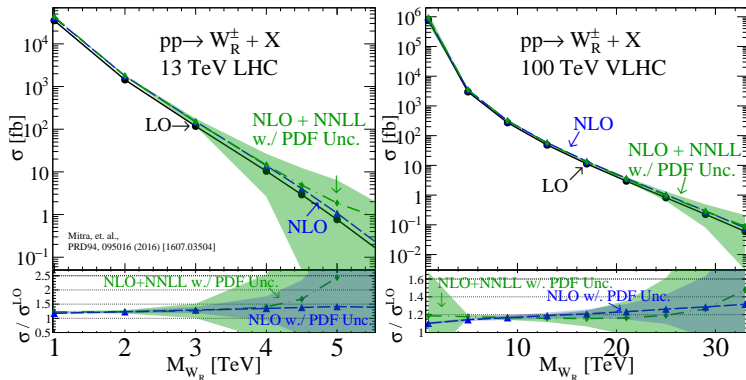
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As  $M_{W_R}^2 \rightarrow s$ , logs diverge since  $M_{W_R}^2 \rightarrow \hat{s} < s$  forces  $g$  to be soft.

In this limit, **soft factorization & exponentiation** possible!

$\Rightarrow$  All-orders (re)summation of  $\alpha_s \log(1 - M^2/\hat{s})$

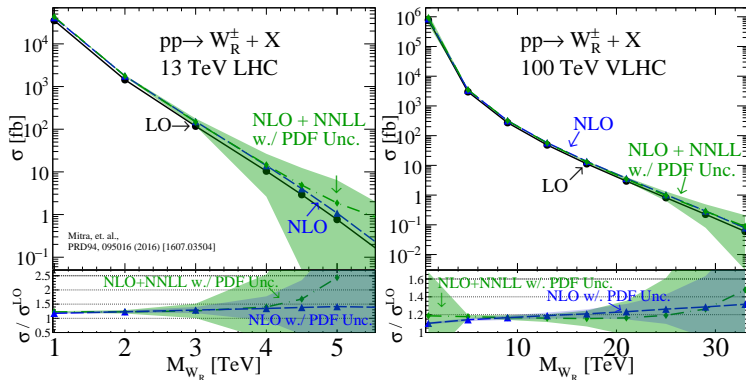
# $W_R$ Numerology at the Edge of Collider Phase Space<sup>10</sup>



At 13 TeV, corrections to production rate  $> +100\%$  for  $M_{W_R} \gtrsim 4.5$  TeV

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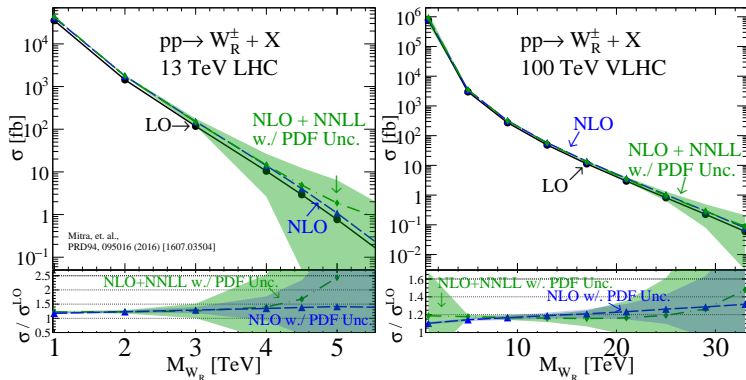
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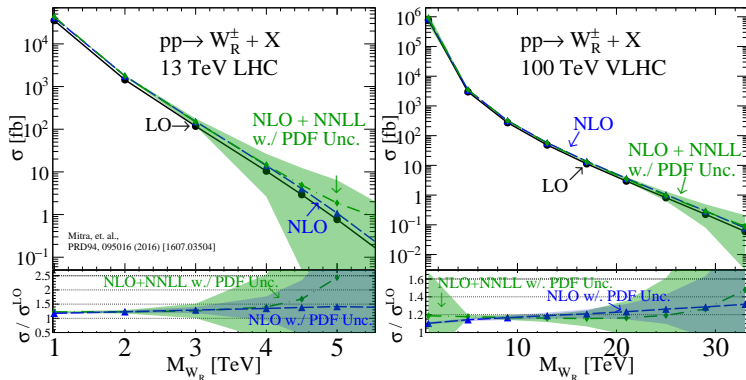


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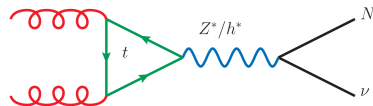
Assuming  $\text{BR} \times \varepsilon \times \mathcal{A} = 2\% \implies N \approx 34 \text{ events } (\sim 6\sigma \text{ vs } \sim 4\sigma)$

<sup>10</sup>Mitra, RR, Scott\*, Spannowsky, PRD ('16) [1607.03504]

**Q: Are high-mass  $W_R$ ,  $Z_R$  unique in this respect?**

# Heavy $N$ at the Edge of Partonic Phase Space<sup>11</sup>

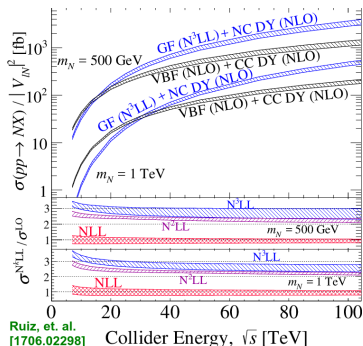
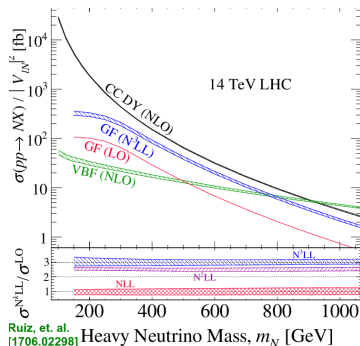
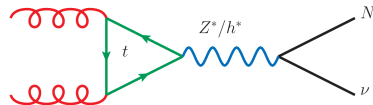
For  $gg \rightarrow N\nu$ , large loops and radiation near  
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 $\Rightarrow$  large increase ( $2 - 3\times$ ) in production rate



<sup>11</sup>Dicus, et al ('85, '91); **TUM** [1408.0983]; **IPPP** [1602.06957, 1706.02298]

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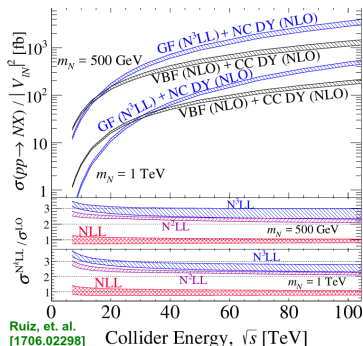
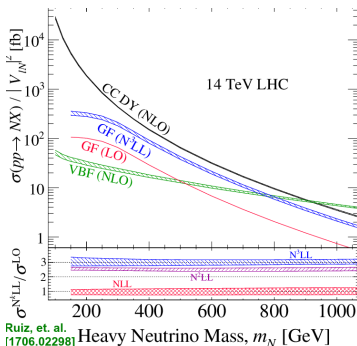
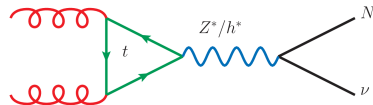
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- At LHC,  $\sigma^{\text{GF}} \sim \sigma^{\text{DY}} \gg \sigma^{\text{VBF}} \Rightarrow \mathcal{O}(1)$  improvement in sensitivity
- For any (proposed) future  $pp$  collider,  $\sigma(N\nu) > \sigma(N\ell)$ !

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## Redux II: Beyond

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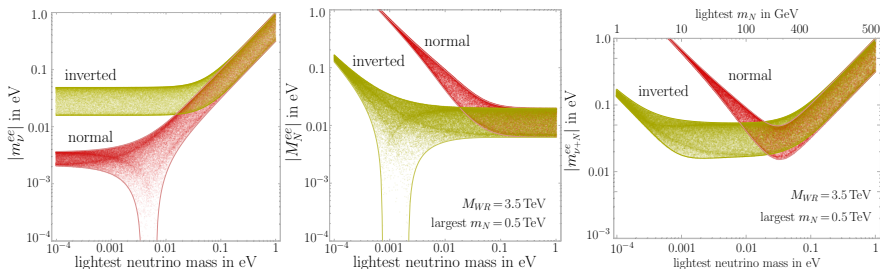
### An outlook of Left-Right Symmetry beyond LHC Run II



# Complementarity to Low Energy Expts

LRSM extremely far reaching in its impact:

- Eg. flavor-changing neutral transitions via Higgs mediation<sup>12</sup>
- Counterexample of naive picture of neutrinoless double beta decay<sup>13</sup>



(L) Canonical description, (C) RH currents, (R) LH+RH currents

<sup>12</sup>Chakraborty, et al [1204.0736]; Bertolini, et al [1403.7112]; Maiezza et al [1407.3678]

<sup>13</sup>E.g., Tello, Nemevsek, Nesti, Senjanovic, Vissani [1011.3522]

- **Immediate:**

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- ▶ Discovery at Run II or elsewhere?
- ▶ **Need:** pheno analyses for “PS boundary” LRSM parameter space
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- ▶ Standardization of pheno tools: adoption of robust, public software

# State-of-Art Event Generators

NLO+PS automated in MadGraph5aMC@NLO, Herwig, Sherpa

- All one needs NLO-accurate FeynRules input model file

**Explosion past two years:** [[feynrules.irmp.ucl.ac.be/wiki/NLOModels](https://feynrules.irmp.ucl.ac.be/wiki/NLOModels)]

- Most neutrino mass models available (just “import” and cite!)

Description	Contact	Reference	FeynRules model files	UFO libraries	Validation material
Dark matter simplified models ( <a href="#">more details</a> )	K. Mawatari	<a href="#">⇒ arXiv:1508.00564</a> , <a href="#">⇒ arXiv: 1508.05327</a> , <a href="#">⇒ arXiv: 1509.05785</a>	-	DMsimp_UFO.2.zip	-
Effective LR symmetric model ( <a href="#">more details</a> )	R. Ruiz	<a href="#">⇒ arXiv:1610.08985</a>	<a href="#">effLRSM.fr</a>	EffLRSM UFO	As of 27 March, updated regularly
GM ( <a href="#">more details</a> )	A. Peterson	<a href="#">⇒ arXiv:1512.01243</a>	-	GM_NLO UFO	
Heavy Neutrino ( <a href="#">more details</a> )	R. Ruiz	<a href="#">⇒ arXiv:1602.06957</a>	<a href="#">heavyN.fr</a>	HeavyN NLO UFO	
Higgs characterisation ( <a href="#">more details</a> )	K. Mawatari	<a href="#">⇒ arXiv:1311.1829</a> , <a href="#">⇒ arXiv:1407.5089</a> , <a href="#">⇒ arXiv: 1504.00611</a>	-	HC_NLO_X0_UFO.zip	-
Inclusive sgluon pair production	B. Fuks	<a href="#">⇒ arXiv:1412.5589</a>	<a href="#">sgluons.fr</a>	<a href="#">sgluons_ufo.tgz</a>	<a href="#">sgluons_validation.pdf</a> ; <a href="#">sgluons_validation_root.tgz</a>
Spin-2 ( <a href="#">more details</a> )	C. Degrande	<a href="#">⇒ http://arxiv.org/abs/1605.09359</a>	<a href="#">dm_s_spin2.fr</a>	SMspin2 NLO UFO	-
Stop pair → t tbar + missing energy	B. Fuks	<a href="#">⇒ arXiv:1412.5589</a>	<a href="#">stop_ttmet.fr</a>	<a href="#">stop_ttmet_ufo.tgz</a>	<a href="#">stop_ttmet_validation.pdf</a> ; <a href="#">stop_ttmet_validation_root.tgz</a>
SUSY-QCD	B. Fuks	<a href="#">⇒ arXiv:1510.00391</a>	-	<a href="#">susyqcd_ufo.tgz</a>	All figures available from the arxiv
Two-Higgs-Doublet Model ( <a href="#">more details</a> )	C. Degrande	<a href="#">⇒ arXiv:1406.3030</a>	-	2HDM_NLO	-
Top FCNC Model ( <a href="#">more details</a> )	C. Zhang	<a href="#">⇒ arXiv:1412.5594</a>	<a href="#">TopEFTFCNC.fr</a>	TopFCNC UFO	-
Vector like quarks	B. Fuks	<a href="#">⇒ arXiv:1610.04622</a>	<a href="#">VLQ_v3.fr</a>	UFO in the 5FNS, UFO in the 4FNS, event generation scripts	All figures available from the arxiv
W/Z' model ( <a href="#">more details</a> )	R. Ruiz, B. Fuks	<a href="#">⇒ arXiv:1701.05263</a>	<a href="#">vPrimeNLO.fr</a>	vPrimeNLO UFO	-

Modern general purpose MC packages are very sophisticated

"With great power there must also come - great responsibility" - S. Lee ('62)

# Outlook

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- **Long-long-term:** Outcome of **near-term** choices. Many discoveries?

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The origin of tiny neutrino masses is still a puzzle and may manifest at **collider experiments** via the production of LRSM partners, e.g.,  $W_R^\pm$ ,  $N$ .

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- Not thrilled, but fact of life and nature
- “Day 1” pheno literature not designed for “edges of phase space”

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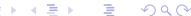
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**Remember:** “The LHC is planned to run over the next 20 years, with several stops scheduled for upgrades and maintenance work.” [press.cern]

- High-Luminosity LHC and Belle II goals:  $1\text{-}5\text{ ab}^{-1}$  and  $50\text{ ab}^{-1}$
- Premature to claim “nightmare scenario” (SM Higgs + nothing else)

<sup>16</sup>Meizza, et al [1503.06834]; Gluza, et al [1604.01388]; **IPPP** [many] 

The logo features a large, light blue oval with a wavy line passing through its center. Inside the oval, the letters 'IP' are written in a large, light blue serif font, with a large '3' to the right. The text 'Thank you.' is centered over the 'IP' in a bold, black sans-serif font.

**Thank you.**